

## ARMY MAP SERVICE TECHNICAL REPORT No. 7 (Rev.)

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# DIRECT AND INVERSE SOLUTIONS OF GEODESICS

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### AMS TECHNICAL REPORT NO. 7 (Revised)

### TABLE OF CONTENTS

			Paragraph	n Page
Section	I.	GENERAL		
		Purpose and Scope	1	. 5
	II.	INTRODUCTION		5
	III.	A RIGOROUS NON-ITERATIVE PROCEDURE FO INVERSE SOLUTION OF VERY LONG GEODES		6
		Preliminary Modification of Helmert's Iterative Solution		6
		Reduction of the Helmert Procedure to Series in x	_	:
		Series in X	• )	<b>;</b>
		Derivation of the Unknown Quantity x	. h	13
		Derivation of the Unknown Quantity x . 4  Determination of Geodetic Distance and Azimuths	13	
		Other Non-Iterative Solutions	. 6	16
		Numerical Illustration of a Sample Solution (International Spheroid) .	. 7	19
		Numerical Coefficients for Other Sphe	eroids 8	21
		Additional Notes on Computational Produces	roce - • • 9	21
		Antipodal Points	10	23
	IV.	TABULAR AND ELECTRONIC COMPUTER METHOFOR NON-ITERATIVE SOLUTION OF GEODET INVERSE, BASED ON SODANO'S PAPER		
		Extension and Modification of Tabula: Electronic Computer Method for Non-I tive Solution of Geodetic Inverse for creased Decimal Accuracy in Short and Lines	tera- r In- i Long	29

	TABLE OF CONTENTS (cont.)		
	Para	graph	Page
v.	TABULAR AND ELECTRONIC COMPUTER METHOD FOR SOLUTION OF DIRECT GEODETIC PROBLEM BASED ON SODANO'S PAPER		38
	Extension of Series of the Direct Geodetic Problem for Greater Accuracy	12	36
Appendix I.	SUMMARY TABLES		
	Part A		39
	Part B		40
Appendix II.	BIBLIOGRAPHY		41

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### ABSTRACT

This Technical Report supersedes TR No. 7 entitled: INVERSE COMPUTATION FOR LONG LINES: A NON-ITERATIVE METHOD BASED ON THE TRUE GEODESIC, which is out of print. It contains the material of the original publication of 1950, and in addition, formulas pertaining to long lines, derived through the years at AMS.

The solutions of the Direct and Inverse Geodetic Problem are presented in forms which are adaptable to desk calculator and to electronic computer.

The maximum errors in the solutions due to the omission of higher order terms have been determined and are presented in tables in the Appendix. These tables will enable the user of the solutions to decide whether the accuracy requirements can be obtained with or without the higher order terms. These higher order terms have been derived and are presented herein.

# DIRECT AND INVERSE SOLUTIONS OF GEODESICS SECTION I. GENERAL

1. Purpose and Scope. The purpose of this report is to present in a single publication the various forms of the Direct and Inverse Solutions of Geodesics which have been solved by the Army Map Service. This report supersedes AMS Technical Report No. 7.

### SECTION II. INTRODUCTION

In Section III of this report a procedure is given for a rigorous and rapid non-iterative inverse solution of very long geodesics.
This procedure, which is in a convenient form for computation by means
of desk calculators, was presented by Mr. Emanuel M. Sodano at the
XIth General Assembly of the International Association of Geodesy and
Geophysics in Toronto, Canada in 1957. The results represent the
gradual extension and accumulated improvements of the original Army
Map Service Technical Report No. 7.

This modification contains a more stable formula for azimuths and an alternative formula for very short lines. More general and accurate formulae for both long and short lines are given herein than are contained in Technical Report No. 7. The complete theoretical derivation starting with a rigorous modification of Helmert's(1) classical formulas are given. The final non-iterative formulas have been extended through terms equivalent to the second, fourth and sixth powers of the eccentricity of the spheroid, and therefore, may be shortened according to the required accuracy.

The solution, which requires no special purpose tables, is accurate to at least the tenth derimal place of radians for the azimuths and the arc distance. If the final formulas are shortened to the second and fourth powers of the eccentricity respectively, the results are accurate to seven and nine decimal places of radians respectively, even for distances circumscribing the earth.

In Section IV the formulas for the solution of the Inverse Geodetic Problem have been adapted to electronic computers. These formulas were derived from the basic formulas of Section III. A solution of the Direct Geodetic problem is given in Section V. The formulas are adapted to electronic computers.

### SECTION III.

A RIGOROUS NON-ITERATIVE PROCEDURE FOR RAPID INVERSE SOLUTION OF VERY LONG GEODESICS

### 2. Preliminary Modification of Helmert's Iterative Solution

e = eccentricity of the spheroid =  $\sqrt{\frac{a_0^2 - b_0^2}{a_0^2}}$ 

e' = second eccentricity =  $\sqrt{\frac{a_0^2 - b_0^2}{b_0^2}}$ 

b<sub>o</sub> = semi-minor axis

L = absolute difference of longitude on the spheroid, between the given endnoints of the geodesic.

 $\beta_1$  and  $\beta_2$  = parametric (or reduced) latitude of the westward and eastward endpoints, respectively.

The relationship between parametric latitude and geodetic latitude is given by the equation  $\tan \beta$  = tan B(l-f) where f is the spheroidal flattening.

\( \) = difference of longitude (approximately L) on the reduced sphere, for which a progressively better value is found with each repetition of the following iteration process:

 $\cos \phi_0 = \sin \beta_1 \sin \beta_2 + \cos \beta_1 \cos \beta_2 \cos \lambda$ 

 $\sin \phi_0 = (\operatorname{sign of } \sin \lambda) \sqrt{1 - \cos^2 \phi_0}$ 

\$\vec{f}\_0 = positive radians

 $\sin 2\vec{p}_0 = 2 \sin \vec{p}_0 \cos \vec{p}_0$ 

 $\sin 3\vec{p}_0 = 3 \sin \vec{p}_0 - 4 \sin^3 \vec{p}_0$ 

 $\cos \beta_0 = (\cos \beta_1 \cos \beta_2 \sin \lambda) \div \sin \overline{\phi_0}$ 

 $\sin^2\!\beta_0 = 1 - \cos^2\!\beta_0$ 

 $\cos 2\sigma = (2 \sin \beta_1 \sin \beta_2 \div \sin^2 \beta_0) - \cos \beta_0$ 

 $\cos 4\sigma = -1 + 2 \cos^2 2\sigma$ 

 $\cos 6\sigma = 4 \cos^3 2\sigma = 3 \cos 2\sigma$ 

 $A' = \frac{e^2e'}{e^1+e} - \frac{e^2e'^2}{16}\sin^2\beta_0 + \frac{3e^2e'^4}{128}\sin^4\beta_0$ 

 $B' = \frac{e^2 e^{12}}{16} \sin^2 \beta_0 - \frac{e^2 e^{14}}{32} \sin^4 \beta_0$ 

 $c' = \frac{e^2 e'^{1}}{256} \sin^{1} \beta_0$ 

 $T = A^{\dagger} \not p_0 = B^{\dagger} \sin \not p_0 \cos 2\sigma + C^{\dagger} \sin 2\not p_0 \cos 4\sigma$ Next approximation to  $\lambda = [(L + T \cos \beta_0) \text{ radians.}]$ 

After a sufficiently accurate  $\lambda$  is found, and using the set

of values from the last iteration, the geodetic distance (S) and arimuths ( $\alpha$ ) between the endooints are obtained as follows:

Finishes (a) between the endocinte are obtained as follows:
$$A_0 = 1 + \frac{e^{\frac{1}{2}}}{4} \sin^2 \beta_0 - \frac{3e^{\frac{1}{4}}}{64} \sinh \beta_0 + \frac{5e^{\frac{1}{6}}}{256} \sinh \beta_0$$

$$P_0 = \frac{e^{\frac{1}{2}}}{4} \sin^2 \beta_0 - \frac{e^{\frac{1}{4}}}{16} \sinh \beta_0 + \frac{15e^{\frac{1}{6}}}{512} \sin^6 \beta_0$$

$$C_0 = \frac{e^{\frac{1}{4}}}{128} \sinh^4 \beta_0 - \frac{3e^{\frac{1}{6}}}{512} \sinh \beta_0$$

$$C_0 = \frac{e^{\frac{1}{6}}}{1636} \sinh^6 \beta_0$$

$$S = b_0 (A_0 \overline{\beta_0} + P_0 \sin \beta_0 \cos 2\sigma - C_0 \sin 2\overline{\beta_0} \cos 4\sigma + D_0 \sin 3\overline{\beta_0} \cos 6\sigma)$$

$$A_{1-2} = \frac{\tan \beta_2 \cos \beta_1 - \cos \lambda \sin \beta_1}{\sin \beta_0}$$

$$\begin{cases} \cot \alpha_{1-2} = \frac{\tan \beta_2 \cos \beta_1 - \cos \lambda \sin \beta_1}{\sin \lambda} \\ \cot \alpha_{2-1} = \frac{\sin \beta_2 \cos \lambda - \cos \beta_2 \tan \beta_1}{\sin \lambda} \end{cases}$$

where  $\varnothing_{1=2}$  and  $\varnothing_{2=1}$  range from  $0^{\circ}$  to  $180^{\circ}$  and  $180^{\circ}$  to  $360^{\circ}$ , respectively, clockwise from north.

### 3. - Andrew Arte tellert Propelure to Gren benier in X

Let it be assumed that the true value of \( \lambda \) is known (that is, the value that would result from an infinite number of Helmert approximations) and let this true value be represented by the given absolute difference of longitude on the spheroid plus a quantity x

which will be determined later.

Thus: 
$$\lambda = (L + x)$$
.

It will be evident, later, that x is a very small positive quantity of the order of  $e^2$ , and therefore well suited for setting up a convergent nower series in x for each expression contained in the Helmert procedure. For example, from the above assumed equation, the following is derived:

$$\cos \lambda = \cos (L + x)$$

$$= \cos L \cos x - \sin L \sin x$$

$$= (\cos L) \left(1 - \frac{x^2}{2} + \dots\right)$$

$$= (\sin L) (x - \dots)$$

Therefore:  $\cos \lambda = (\cos L) - (\sin L) x - \frac{1}{2}(\cos L) x^2 + \dots$ 

There is thus available, at the outset, a series for the true  $\cos \lambda$  with which to begin the Helmert solution and develop it in power series in x in its entirety. The process consists of substituting each new series into the succeeding Helmert expressions as required. For convenience, the following additional notation will be used:

$$N = e' \stackrel{?}{=} (e' + e)$$

$$a = \sin \beta_1 \sin \beta_2$$

$$b = \cos \beta_1 \cos \beta_2$$

$$\cos \vec{p} = a + b \cos L$$

$$c = b \sin L \csc \vec{p}$$

$$m = 1 - c^2$$

$$h = e^{i2m}$$

$$V_1 = (\tan \beta_2 \cos \beta_1 - \cos L \sin \beta_1) \cdot \sin L$$

$$V_2 = (\sin\beta_2 \cos L - \cos\beta_2 \tan\beta_1) = \sin L$$

Listed, below, in the same sequence as the corresponding Helmert expressions, is the complete set of series through  $\mathbf{r}\cos\boldsymbol{\beta}_{\bullet}$ . The extent of the powers of x is such as to permit accuracies of the  $\mathbf{e}^{\lambda}$  order in  $\lambda$ , for subsequent application to the distance and azimuths to the same degree of accuracy as the reference Helmert iteration form.

cos 
$$f_0$$
 =  $(\cos \phi)$  -  $(\cos \phi)$  x -  $\frac{1}{2}(e^2 \cos \phi + P \sin \phi)$  x<sup>2</sup>  
con  $f_0$  =  $(\sin \phi)$  +  $(\cos \phi)$ x -  $\frac{1}{2}(e^2 \sin \phi - P \cos \phi)$  x<sup>2</sup>  

$$f_0$$
 =  $\phi$  +  $(f_0 - \phi)$  =  $\phi$  + are  $\sin (\sin (f_0 - \phi))$   
=  $\phi$  + are  $\sin (\sin \phi)$  cos  $\phi$  -  $\cos \phi$  sin  $\phi$ )  
=  $\phi$  +  $(\cos x + \frac{1}{2}(P) + x^2)$   
sin  $2f_0$  =  $2 \sin \phi \cos \phi$ .  
sin  $\lambda$  =  $(\sin 1)$  +  $(\cos 1)$  x -  $\phi$  sin 1) x<sup>2</sup>  
cos  $\beta$  =  $(e)$  +  $(e)$  x -  $\frac{1}{2}(em + 3e P \cot \phi)$  x<sup>2</sup>  
cos  $\beta$  =  $(e)$  +  $(e)$  x -  $(em + 3e P \cot \phi)$  x<sup>2</sup>  
cos  $\beta$  =  $(em + (em + 3e P \cot \phi)$  x =  $(em +$ 

-  $4cP^2 \sin \emptyset x$ 

-B' 
$$\sin \phi_0 \cos 2\sigma = \frac{e^2}{128}$$
 (- 8h  $\sin \phi \cos \phi + 16e^{-2} P \sin^2 \phi$   
+  $\ln^2 \sin \phi \cos \phi - 8e^{-2} hP \sin^2 \phi$ )  
-  $\frac{e^2}{16}$  (hc) x

C' 
$$\sin 2 \vec{\phi}_0 \cos 4 \vec{\phi} = \frac{e^2}{128} \left( h^2 \sin \vec{\phi} \cos \vec{\phi} - 2h^2 \sin^3 \vec{\phi} \cos \vec{\phi} \right)$$

$$- 8e^{i^2} h^2 \sin^2 \vec{\phi} \cos^2 \vec{\phi} + 8e^{i4} \vec{\phi}^2 \sin^3 \vec{\phi} \cos \vec{\phi}$$

$$\Upsilon\cos\beta_0 = \frac{e^2}{128} \left(128\text{Nc} - 8\text{hc} - 8\text{hc} \sin \phi \cos \phi + 16e^{12}\text{cP} \sin^2 \phi + 3\text{h}^2\text{c}\phi + 5\text{h}^2\text{c} \sin \phi \cos \phi - 2\text{h}^2\text{c} \sin^3 \phi \cos \phi \right)$$

$$- 8e^{12}\text{hcP} \sin^2 \phi - 8e^{2}\text{hcP} \sin^2 \phi \cos^2 \phi$$

$$+ 8e^{14}\text{cP}^2 \sin^3 \phi \cos \phi + \frac{e^2}{16} \left(16\text{Nc}^2 + 16\text{NP}\phi - 2\text{hc}^2 + 16\text{NP}\phi + 2e^{12}\text{c}^2 + 2$$

### L. Derivation of the Unknown Quantity x

Since the substitution into the Helmert iteration began with an alrebraic series representing the true  $\lambda$ , the next approximation to  $\lambda$  must of necessity be its equal; that is:

The next approximation to  $\lambda$  = the starting true  $\lambda$ 

or 
$$L + \Upsilon \cos \beta_0 = L + x$$

and therefore 
$$\Upsilon \cos \beta_0 = x$$
.

By replacing  $T\cos\beta_0$  with its corresponding power series, the above equation takes the following quadratic form:

$$Q_1 + Q_2 x + Q_3 x^2 = x$$

for which the required solution of x to the proper order is

$$x = Q_1(1 + Q_2 + Q_2^2 + Q_1Q_3).$$

Finally, substituting for  $Q_1$ ,  $Q_2$  and  $Q_3$ , produces the following end result:

$$x = \frac{e^2c}{128} \int 128N \vec{p} + 128e^2N^2c^2\vec{p} - 8h\vec{p} - 8h \cdot \sin \vec{p} \cos \vec{p} + 128e^2N^2F \vec{p}^2$$

- +  $16e^{12}P \sin^2\theta + 128e^{\ln 3}c^{\ln \theta} 2\ln^2 Nhc^2\theta + 3h^2\theta$
- $8e^2Nhc^2 \sin \varphi \cos \varphi + 5h^2 \sin \varphi \cos \varphi 6he^h A^3c^2 m$
- 2h<sup>2</sup> sin<sup>3</sup> cos + (16e<sup>2</sup>e<sup>2</sup>N + lik8e<sup>4</sup>N<sup>3</sup>) c<sup>2</sup>P 2
- 16e<sup>2</sup>NhP ¶ + 16e<sup>2</sup>e '2Nc<sup>2</sup>P sin<sup>2</sup> ¶ 8e '<sup>2</sup>hP sin<sup>2</sup> ¶
- 16e2NhP & sin & cos & 1904N3c2P \$3 cot \$
- 8e'2/F sin2 \$\int \cos2 \$\int\$ + 128e 4 3p2 \$\int\$
- + 32e<sup>2</sup>e<sup>12</sup>HP<sup>2</sup>  $\int \sin^2 \int d + 8e^{-1} \ln^2 \sin^3 \int d \cos \int d$

The above rigorously developed expression is completely non-iterative, since it requires only the given spheroidal longitude. It therefore permits a direct evaluation of the ultimately true  $\lambda$  (that is, L + x), extended in this case through terms equivalent to the  $e^2$ ,  $e^4$  and  $e^6$  order consecutively, in accordance to the accuracy that may be desired. Furthermore, it represents the algebraic solution of the hitherto unknown quantity x used in the power series version of each of the intermediate Helmert expressions.

### 5. Determination of Geodetic Distance and Azimuths

The non-iterative expression that has been developed for  $\mathbf{x}$  suggests at once a numerical solution of distance and azimuths wherein, using the resulting true value of  $\lambda$ , only a single evaluation of Helmert's original formulas is necessary. An illustrative example by such a procedure is given in **paragraph**  $\P$ .

On the other hand, instead of reverting to functions of the true  $\lambda$ , the distance and azimuths themselves can be expanded non-iteratively into power series of x with coefficients in terms of the given spheroidal difference of longitude. This is accomplished below, but limited to the  $e^{i\mu}$  order of accuracy, since this manner of obtaining the distance and azimuths through  $e^6$  would require each component series to one higher power of x than was necessary for  $\lambda$ . Again, the series are developed in the same sequence as

the corresponding Helmert expression.

$$A_{0} = \frac{1}{64} (64 + 16h - 3h^{2}) - \frac{1}{2} (e^{12}cP) \times$$

$$E_{0} = \frac{1}{16} (4h - h^{2}) - \frac{1}{2} (e^{12}cP) \times$$

$$C_{0} = \frac{h^{2}}{128}$$

$$A_{0} = \frac{1}{64} (64 + 16h - 3h^{2}) + \frac{1}{4} (4c + hc - 2e^{12}cP) \times$$

$$+ \frac{1}{2} (P) x^{2}$$

 $P_{c}\sin \mathbf{f}_{o}\cos 2\mathbf{\sigma} = \frac{1}{6h} (16h \sin \mathbf{f} \cos \mathbf{f} - 32e^{i2p} \sin^{2}\mathbf{f} - 4h^{2}\sin \mathbf{f} \cos \mathbf{f} + 8e^{i2h}P \sin^{2}\mathbf{f}) + \frac{1}{4} (hc) x$ 

$$S = \begin{bmatrix} \frac{b_0}{6U} & (6U\vec{p} + 16h\vec{p} + 16h \sin \vec{p} \cos \vec{p} - 32e^{i2}P \sin^2\vec{p} \\ -3h^2\vec{p} - 5h^2ain & (6cca \vec{p} + 3)^2cin & cos^2\vec{p} \\ + 8e^{i2}hP \sin^2\vec{p} + 8e^{i2}hP \sin^2\vec{p} \cos^2\vec{p} \\ - 8e^{i4}P^2 \sin^3\vec{p} \cos^2\vec{p} + \frac{b_0}{2} (2c + hc - e^{i2}cF\vec{p}) \times \frac{b_0}{2} (P) \times^2 \end{bmatrix}$$

$$\cot \alpha_{1-2} = \left[ v_1 - \left( \frac{v_2 \cos \beta_1}{\sin L \cos \beta_2} \right) x + \left( \frac{v_1}{2 \sin^2 L} + \frac{v_2 \cos L \cos \beta_1}{2 \sin^2 L \cos \beta_2} \right) x^2 \right]$$

The x and  $x^2$  for the above formulas of distance and azimuths can be substituted either numerically or algebraically using, in this case, only the first 6 terms of x for accuracies equivalent to the  $e^{\frac{1}{4}}$  order. The <u>algebraic</u> substitution gives the following final expressions:

$$S = \frac{b_0}{6l_4} \left\{ 6l_4 f^2 + 6l_4 e^2 N c^2 f^2 + 16h_4 f^2 + 16h \sin f \cos f \right.$$

$$- 32e^{12}P \sin^2 f^2 + 6l_4 e^{1}N^2 c^{1} f^2 - 3h^2 f^2 + (32e^2N - 4e^2)hc^2 f^4 - 4e^2 hc^2 \sin f \cos f + 2h^2 \sin^3 f \cos f + (96e^{1}N^2 - 32e^2 e^{12}N)c^2 P f^2 + 8e^2 e^{12}c^2 P \sin^2 f + 8e^{12}hP \sin^2 f + 8e^{12}hP \sin^2 f \cos^2 f - 8e^{1} hp^2 \sin^3 f \cos f \right\}$$

$$\cot \mathcal{A}_{1-2} = U_1 - \frac{e^2 N c f U_2 \cos f f_1}{\sin L \cos f f_2} - \frac{e^{1}N^2 c^3 f U_2 \cos f f_1}{\sin L \cos f f_2}$$

$$+ \frac{e^2 h c f U_2 \cos f f_1}{16 \sin L \cos f f_2} + \frac{e^2 h c U_2 \sin f \cos f \cos f f_1}{16 \sin L \cos f f_2}$$

$$- \frac{e^1 N^2 c P f^2 U_2 \cos f f_1}{\sin L \cos f f_2} - \frac{e^2 e^{12} c P U_2 \sin^2 f \cos f f_1}{8 \sin L \cos f f_2}$$

$$+ \frac{e^1 N^2 c^2 f^2 U_2 \cos L \cos f f_1}{2 \sin^2 L \cos f f_2} + \frac{e^1 N^2 c^2 f^2 U_1}{2 \sin^2 L}$$

The corresponding  $\cot \alpha_{2-1}$  is obtainable from the above by interchanging  $v_1$  with  $v_2$  and  $\beta_1$  with  $\beta_2$ .

Thus, progressively, there have been develored three rigorous methods for determining geodetic distance and azimuths non-iteratively: as a function of the true  $\lambda$ , as a power series in x, and culminated by an <u>explicit</u> expression in essentially the given spheroidal latitude and longitude of the endpoints. For shorter lines, or for reduced accuracy on long lines, terms may be still further eliminated according to the next higher powers of  $e^2$ ,  $e^{12}$ , h and x, or equivalent combinations thereof.

### 6. Other Non-Iterative Solutions

The distance and azimuths by the original Helmert method are essentially functions of elements in the following spherical triangle:

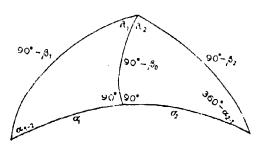


Fig. 1.

where

$$\lambda = \lambda_2 - \lambda_1$$

$$\phi_0 = \sigma_2 - \sigma_1$$

$$2\sigma = \sigma_2 + \sigma_1$$

and  $\lambda_1$ ,  $\lambda_2$ ,  $\sigma_1$  and  $\sigma_2$  are regarded as negative or positive according to whether they are west or east of the perpendicular arc  $90^c$ - $\beta_c$ .

(For this specific configuration, therefore,  $\lambda$  and  $\oint_0$  actually represent the sum of the absolute components and  $2\sigma$  the difference.)

Since Helmert's method of successive approximations can only determine  $\lambda$  first, the subsequent solution of the above spherical triangle would always begin with  $\lambda$  and the known  $A_1$  and  $A_2$ . The present paper, however, has developed not only a non-iterative expression for  $\lambda$ , but also independent noner series for the various elements of this spherical triangle or functions thereof. Therefore the combination of ways to compute quantities leading to the distance and azimuths is increased considerably. In addition, the x for such series can be substituted either numerically or algebraically, in the manner shown for the distance and azimuth series in

The above notentiality for increasing the number of non-iterative solutions may be seen from the expressions (1) below, where-in the x and  $x^2$  of the  $\mathbf{p}_0$  series are algebraically eliminated.

$$\begin{cases}
\oint_{0} = \oint_{0}^{2} + \frac{16e^{2N-e^{2}e^{\frac{12}{2}}}}{16} (c^{2}\oint_{0}^{2}) + \frac{16e^{4N^{2}+e^{2}e^{\frac{12}{2}}}}{16} (c^{4}\oint_{0}^{2}) \\
+ \frac{e^{2}e^{\frac{12}{2}}}{16} (c^{2}\sin\oint_{0}^{2}\cos\oint_{0}^{2}) - \frac{e^{2}e^{\frac{12}{2}}}{16} (c^{4}\oint_{0}^{2}\cos\oint_{0}^{2}) - \frac{3e^{4N^{2}}}{2} (c^{4}\oint_{0}^{2}\cos\oint_{0}^{2})
\end{cases} (1)$$

$$- \frac{e^{2}e^{\frac{12}{2}}}{8} (ac^{2}\sin\oint_{0}^{2}) + \frac{3e^{4N^{2}}}{2} (c^{2}d\oint_{0}^{2}) - \frac{3e^{4N^{2}}}{2} (c^{4}\oint_{0}^{2}\cot\oint_{0}^{2})$$

The computed value of  $f_0$  is then combined with  $\mathcal{B}_1$  and  $\mathcal{B}_2$  to obtain  $\boldsymbol{x}$ 's, followed by  $\mathcal{B}_0$ ,  $2\boldsymbol{\sigma}$ ,  $A_0$ ,  $B_0$ ,  $C_0$  and finally the geodetic distance. When adopting such varied procedures for solving

the reference triangle, care should be taken to aveid formulations which lead to a weak determination of required quantities. These difficulties may most likely occur at extremes of latitude, lorgie tude, or azimuth.

The non-iterative series, too, are functions of elements of a spherical triangle, but defined by  $eta_1$  and  $eta_2$  and the given longitude L. This amounts simply to a substitution of L for  $\lambda$ , which results in a spherical triangle with parts corresponding as follows:

Series  $\beta_1$   $\beta_2$  L  $\bar{p}$  c  $v_1$   $v_2$ Helmert  $\beta_1$   $\beta_2$   $\lambda$   $p_0$   $\cos \beta_0$   $\cot \alpha_{1-2}$   $\cot \alpha_{2-1}$ Starting with the given  $\beta_1$ ,  $\beta_2$ , and L, the values of all quantities used in the non-iterative series may thus be solved trigonometrically in various orders.

It is also to be noted that in the relation  $\lambda = (L + x)$ , if x is assumed to be zero, L is considered to be equal to  $\lambda$ . Therefore in the various nower series in x, the <u>constant term</u> can represent the true value of the series by simply replacing functions of L with  $\lambda$ . This is well illustrated by the  $A_0$  series in x in recognity  $\xi$  and its counterpart in parameter  $\xi$  is the empty arrange. The principle can well be incorporated in computation forms, such as the one on the next  $\xi$  applied to  $\xi$ .

### 7. Numerical Illustration of a Sample Solution (International Spheroid)

```
x_{red} = A [ \mathcal{I}(237.2388918 + B) + \sin \mathcal{I}(C+D) 
                                                              0.00326 58167
                  +C(F+E) } + 70519.51145
         ) = 1. + x
                                                          106911113".6230"
                                                              0.96035 63900
      \min \lambda
                                                             -0.27877 518148
      cos
  \cos \tilde{\phi}_0 = a + b \cos \lambda
                                                              0.05509 74283
  \sin \mathcal{L}_0 = (\text{sign of } \sin \lambda) \sqrt{1 - \cos^2 \theta_0}
                                                             0.99848 09830
       تر = rositive radians
                                                             1.51567 09835
 \sin 2f = (\sin f_0 \cos f_0) \div 0.5
                                                             0.11002 71687
   \cos \beta_c = (b \sin \lambda) + \sin \phi_c
                                                             0.64012 07839
          g = 1 - \cos^2 \beta_0
                                                              0.58986 12195
   ..... 2 σ - (28 - q cos Øn) + q
                                                             0.76108 89231
   cot \# \sigma = (cos^2 2\sigma - 0.5) + 0.5
                                                              0.15851 26977
         H = 6356911.946 + 10756.165q =
                                                      6363251.841
         1 = 10756.165a - 18.200a^2
                                                          6338.312
        J = 2.2750^2
                                                              0.792
 -id_{j-1} = (\tan \beta_2 \cos \beta_1 - \cos \lambda \sin \beta_1)
                                                              1.07455 96397
   \alpha_{0.1} = (\sin \beta_2 \cos \lambda - \cos \beta_2 \tan \beta_1) -0.47245 22891
    Clockwise from North, in quad
I or II for cot + or -, resp.
                                                        42056130".03557
    \alpha_{/-1} = Clockwise from North, in quad
III or IV for cot + or -, resp. 295017*18".59765
numericus = H_{O}^{\phi} + I \sin \phi_{O} cos2 \sigma -
                                                              9649412.85Lm.
                    J sin 22 cosh o
```

### 8. Numerical Coefficients For Other Spheroids

The illustrative solution given in the preceding section contains fixed numerical coefficients which are functions solely of the size and shape of the International spheroid. The algebraic expressions of these coefficients, together with their values, are shown below in the order of appearance in the sample solution. For any other spheroid, these expressions can be quickly re-evaluated once and for all and substituted for the corresponding

International values. (Note: e2N = flattening.)  $+\sqrt{1-e^2}$ 0.99663 29966  $(16e^2N^2 + e^{1^2}) \div e^{1^2} =$ 4.9865 20649  $2e^{12} \div (16e^2N^2 + e^{12}) =$ 0.40108 12630  $16e^2N^2 = (16e^2N^2 + e^{12}) =$ 0.79945 93686  $16e^2N^2 \div e^{i^2} =$ 3.9865 20649  $(16N - e^{1^2}) = (16e^2N^2 + e^{1^2}) =$ 237.2388 918  $16 \div e^2(16e^2N^2 + e^{*2}) = 70519.51145$ 

 $b_0 = 6356911.946$   $b_0 e^{12} \cdot 4 = 10756.165$   $3b_0 e^{14} \cdot 64 = 13.650$   $b_0 e^{14} \cdot 16 = 18.200$   $b_0 e^{14} \cdot 128 = 2.275$ 

### 9. Additional Notes on Computational Procedures

Although the illustrative solution given in production of the prod

is primarily intended for accuracy equivalent to the e<sup>ll</sup> order, it easily lends itself to any required degree. This is accomplished simply by adding or subtracting appropriate terms of x, H, 1, J, and S. The extended terms are given in the latter part of Section III, paragraphs leand? I, respectively. For short lines or reduced accuracy on long lines, x on the International spheroid becomes merely (AP + 297) and all terms in q<sup>2</sup> are omitted, with the consequent elimination of many other supporting quantities. Similar savings are realized for other forms of solutions presented herein.

For short lines, the resulting small  $\not \! D$  is computed more accurately from  $\sin \vec p \!\!\! D$  obtained as follows:

$$\sin \frac{\sqrt{2}}{2} = + \sqrt{b} \sin^2 \frac{L}{2} + \sin^2 \frac{B_1 - B_2}{2}$$

$$\cos \frac{\sqrt{2}}{2} = (\text{sign of sinL}) \sqrt{0.5(1 + \cos \sqrt{2})}$$

$$\sin \mathcal{J} = \left(\sin \frac{\mathcal{J}}{2} \cos \frac{\mathcal{J}}{2}\right) \stackrel{2}{\sim} 0.5$$

Similarly,  $\sin \oint_0$  is obtained as above by replacing  $\oint$  with  $\oint_0$  and L with  $\lambda$ . In either case, squaring the small sines under the radical increases their significant decimal places.

If the numerator of x is to be cumulated in a ten digit calculator, 9 decimal places should be allotted to  $\mathcal{G}$ , sin  $\mathcal{G}$  and  $\mathcal{G}$ , but only 7 decimals to their multipliers. However, when the value

of G is 10 or greater, decrease its decimal places accordingly and increase those of F and E correspondingly. For a smaller calculator, reduce all decimals equally.

Use co-function of  $anoldsymbol{eta}$  or  $\cot \sigma$  when their values are too large.

Thus 
$$\cot \beta_n = \frac{\cot B_n}{+\sqrt{1-e^2}}$$
 and  $\tan \alpha = \frac{1}{\cot \alpha}$ 

The accuracy of geodetic distances computed through the  $e^2$ ,  $e^4$  and  $e^6$  order for <u>very long geodesics</u> is within a few meters, centimeters and tenths of millimeters respectively. Azimuths are good to tenths, thousandths, and hundred thousandths of a second. Further improvement of results occurs for shorter lines.

Some of the terms in the sample solution of paragraph 7 have been grouped for ease of computing by desk calculator. For electronic computers, however, the terms are best left in series form, thus being ideally suited to adding or removing them according to accuracy requirements.

### 10. Antipodal Points

In the various series that have been presented,  $\not p$  represents a spherical arc distance which varies from 0° to 180° and even to 360° according to whether the recodetic line is very short, half around the earth or completely around it. At these specific instances, quantities such as  $\csc \not p$ ,  $\cot \not p$ , and P approach infinity. For the case of the very short lines, this condition is equalized

Take factors of and sin of which crossed by some. For the cross the course we have been some form the formal and the formal argument.

Closer inspection of the various series in x shows, nevertheless, that this condition of divergence never provails in the constant corns, and for succeeding coefficients it is to me crosses fugges when the power of the corresponding x. Therefore, have not it could be equalized if x were sufficiently guall.

The first equation of the transfer as follows:

$$\lambda = (L + x).$$

this true value of A could have been represented, instead, by:

$$\lambda = (\mathbf{I}_n + \mathbf{y})$$

where In is an orbitrary amount or lored but a type meanly equal to  $\lambda$  one therefore  $\chi$  is correspondingly smaller than x. This new assume the saids to a cot of newer servers in  $\lambda$  south and identical to the end, except that its coefficients of the bard facetion of  $\lambda$ , which is the obvious value to assign to  $\lambda_1$  south be the suightly that there except of solving as employees and in occurs of the lores

The relation derived at the hopfishing of the real land will mercingly charge from:

where, as noted, the substitution of the  $T\cos\beta_0$  series given at the end of paragraph 3—will now be in terms of  $L_n$  and x instead of L and x. Solving the above equation for z (this time through only the  $e^{\frac{1}{4}}$  order of accuracy) gives:

$$x = \frac{16(I-I_n) + (16e^2Nc\sqrt{9} - e^2hc\sqrt{9} - e^2hc\sin\sqrt{9}\cos\sqrt{9} + 2e^2e^{\sqrt{2}}cP\sin^2\sqrt{9})_n}{16(1 - e^2Nc^2 - e^2NP\sqrt{9})_n}$$

where the subscripts in to the parenthesis indicate that c,  $\vec{A}_{i}$  h, P, etc. are functions of  $L_{i}$  instead of L. This time, the denominator of the expression cannot be algebraically divided into the numerator, because the  $e^{2}NP\vec{p}$  term is relatively large for nearly antipodal lines.

With the above correction z to an arbitrary but sufficiently accurate value  $L_n$ , the true  $\lambda$  of antipodal lines is essentially obtained again non-iteratively, and therefore more rapidly than by numerous individual successive approximations. Thus, also, a previous  $h^n$  longitude discrepancy noted by Mr. H. F. Rainsford for a line of about 179°46'18" longitude would be resolved. In this connection, appreciation is expressed to Mr. Rainsford for his interest in the subject which resulted in profitable correspondence.

### SECTION IV

TAPULAR AND ELECTRONIC COMPUTER METHOD FOR NON-ITERATIVE SOLUTION OF GEODETIC INVERSE, BASED ON CODANO'S PAPER

Due to their series-like nature, the formulas given in this section for distance and azimuth are more adaptable to electronic computer

programming than the corresponding closed formulas of Section III.

This method (unlike the one just discussed) does not have the restriction that P<sub>1</sub> and L<sub>1</sub> <u>must</u> be the latitude and longitude, respectively, of the westward point. Here, B<sub>1</sub> and L<sub>1</sub> are the geometric latitude and longitude, respectively, of any point.

The distance equation of this section was derived by making the Following substitutions into the distance S equation on page 1/4:

$$f = e^{2}N$$

$$e^{t^{2}m} = h \text{ (where } e^{t^{2}} \text{ was expressed in terms of } f)$$

$$m = 1 - c^{2}$$

$$P = (1 - c^{2}) \cot \emptyset - a \csc \emptyset$$

The expression ( $\lambda$  - L) of this method is equivalent to "x" on page 10. The series for ( $\lambda$  - L) was derived by making the substitutions (1) into the equation for "x" on page 11. The computation for this method is as follows:

Pl, Li = geographic latitude and longitude, respectively, of any point

 $h_2$ ,  $h_2$  = geographic latitude and longitude, respectively, of any other point, non-antipodal

Latitudes and longitudes considered (4) month and east, (-) south and west

Legained: Q, S = azimuths clockwise from north and distance between points, respectively.

$$\frac{S}{b_0} = \left[ (1 + f + f^2) \vec{p} \right] 
+ a \left[ (f + f^2) \sin \vec{p} - \frac{f^2}{2} \vec{p}^2 \csc \vec{p} \right] 
+ m \left[ - (\frac{f + f^2}{2}) \vec{p} - (\frac{f + f^2}{2}) \sin \vec{p} \cos \vec{p} + \frac{f^2}{2} \vec{p}^2 \cot \vec{p} \right] 
+ a^2 \left[ - \frac{f^2}{2} \sin \vec{p} \cos \vec{p} \right] 
+ m^2 \left[ (\frac{f^2}{16}) \vec{p} + \frac{f^2}{16} \sin \vec{p} \cos \vec{p} - \frac{f^2}{2} \vec{p}^2 \cot \vec{p} - \frac{f^2}{6} \sin \vec{p} \cos^3 \vec{p} \right] 
+ am \left[ (\frac{f^2}{2}) \vec{p}^2 \csc \vec{p} + \frac{f^2}{2} \sin \vec{p} \cos^2 \vec{p} \right]$$

$$\frac{\left(\frac{\lambda - 1}{\sqrt{c}}\right)}{\sqrt{c}} = \left[ (f + f^2) \sqrt{d} \right] + a \left[ -\left(\frac{f^2}{2}\right) \sin \sqrt{d} - f^2 \sqrt{d} \cos \sqrt{d} \right]$$

$$= r \left[ -\frac{5f^2}{4} \sqrt{d} + \frac{f^2}{4} \sin \sqrt{d} \cos \sqrt{d} + f^2 \sqrt{d} \cos \sqrt{d} \right]$$

where:  $a_0$ ,  $b_0$  = semi-major and semi-minor axes, respectively, of spheroid

f = spheroidal flattening =  $(1 - \frac{b_0}{a_0})$ 

P = number of seconds in one radian = 206,264.80625

 $L = (L_2-L_1)$  or  $(L_2-L_1) + \{\text{sign opposite of } (L_2-L_1)\} 360^\circ$ 

The plinter L has an absolute value or >1800; according to whether the shorter or longer geodetic arc is required. However, for meridional arcs ( $|L| = 0^{\circ}$  or  $180^{\circ}$  or  $360^{\circ}$ ) use either L but consider it as (+) for the shorter and (-) for the longer arc.

> $tan \beta = tan B (1-f) when |B| \le 450 \text{ or } cot \beta = \frac{\cot B}{1-f}$  when (B) 450

 $a = \sin \beta_1 \sin \beta_2$ 

 $h = \cos \beta_1 \cos \beta_2$ 

cos  $\vec{d} = a + b \cos L$ sin  $\vec{p} = \pm \sqrt{(\sin L \cos \beta_2)^2 + (\sin \beta_2 \cos \beta_1 - \sin \beta_1 \cos \beta_2 \cos L)^2}$ 

The sign of  $\sin \phi$  is (+) or (-) according to whether the shorter or the longer arc is required. The quantity under the radical and its root must be computed by floating decimal to obtain  $ec{g}$  to full accuracy for short lines.

> f = resitive radians (obtain reference angle from gor  $\cos { ilde{m{g}}}$  which we have smaller absolute value.)

c = (b sir. L) = sin Ø

 $\mathbf{r}_{i} = \mathbf{1} - \mathbf{3}^{2}$ 

$$\cot \alpha_{1-2} = (\tan \beta_2 \cos \beta_1 - \cos \lambda \sin \beta_1) \div \sin \lambda$$
 Cmit for 
$$\cot \alpha_{2-1} = (\cos \lambda \sin \beta_2 - \tan \beta_1 \cos \beta_2) \div \sin \lambda$$
 arcs
If  $|\cot \alpha| > 1$ , divide result into 1 to obtain  $\tan \alpha$  instead.

		Quadrant of 1-2		Quadrant of
Sign	Sign of tan or (cot) 1-2		Sign of tan or (cot)	
+	+	I	ŧ.	111
	-	11	-	2.0
-	+	111	,	ī
	. <b>-</b>	W	_	11

For meridional area, enter the above table with the sign of the numerator of cot  $\boldsymbol{\bowtie}$ , and reference and le  $\theta^0$ .

11. Extension and Modification of Tabular & Electronic Computer Relact for Mos-Itanualna Salution of Modific Inventor For Incressed Decidal Scourage in Sacra and Lour Lines

and its constraint the distance of what we observe by addition that  $\hat{E}_{i}$  becomes to the  $\frac{\partial}{E_{i}}$  coming them there. The series that vectors:

$$\frac{S}{b_0} = (1+f+f^2+f^3) \int_{0}^{\infty} \int_{0}^{\infty$$

Similarly, the accuracy of the  $\frac{(\lambda - 1)}{Rc}$  series may be extended by adding the  $f^3$  term. The series  $t^3$  en becomes:

$$\frac{\lambda_{-1}}{P^{c}} = \left[ (f + f^{2} + f^{3}) \vec{f} \right] + a \left[ (-\frac{1}{2}f^{2} - f^{3}) \sin \vec{f} + (-f^{2} - Lf^{3}) \vec{g}^{2} \csc \vec{f} \right]$$

$$+ \frac{3}{2} f^{3} \vec{f}^{3} \csc \vec{f} \cot \vec{f} \right] + m \left[ (-\frac{5}{L} f^{2} - 3f^{3}) \vec{g} + (\frac{1}{2}f^{2} + \frac{1}{2}f^{3}) \sin \vec{f} \cos \vec{f} \right]$$

$$+ (f^{2} + Lf^{3}) \vec{f}^{2} \cot \vec{f} - \frac{1}{2}f^{3} \vec{f}^{3} - \frac{3}{2} f^{3} \vec{f}^{3} \cot^{2} \vec{f} \right]$$

$$+ m^{2} \left[ \frac{31}{16} f^{3} \vec{f} - \frac{7}{16}f^{3} \sin \vec{f} \cos \vec{f} + \frac{1}{2}f^{3} \vec{f}^{3} - \frac{1}{8}f^{3} \sin^{3} \vec{f} \cos \vec{f} \right]$$

$$- \frac{2}{2}f^{3} \vec{f}^{2} \cot \vec{f} + \frac{1}{2}f^{3} \vec{f} \cos^{2} \vec{f} + \frac{5}{2} f^{3} \vec{f}^{3} \csc \vec{f} \cot^{2} \vec{f} - \frac{7}{2} \sin \vec{f} \cos^{2} \vec{f} \right]$$

$$+ am \left[ \frac{9}{2}f^{3} \vec{f}^{2} \csc \vec{f} - \frac{3}{2}f^{3} \vec{f} \cos \vec{f} - \frac{7}{2}f^{3} \vec{f}^{3} \csc \vec{f} \cot \vec{f} - \frac{f^{3}}{2} \sin \vec{f} \cos^{2} \vec{f} \right]$$

$$+ f^{3} \sin \vec{f} + a^{2} \left[ f^{3} \vec{f} + \frac{1}{2}f^{3} \sin \vec{f} \cos \vec{f} + f^{3} \vec{f}^{3} \csc^{2} \vec{f} \right]$$

The  $f^3$  term of the above series has been maximized and this value is shown in Appendix I, Part A. The error in the azimuth  $\varphi$  which would result from the omission of the  $f^3$  term of  $\frac{(\lambda - L)}{f^2c}$  has been recorded in Appendix I. Part A.

In the case of short geodetic lines (lines shorter than 180 miles) when the values of  $\vec{p}$ ,  $\lambda$ -L, etc. of the series above become small, it is necessary to use floating point notation in order to insure greater decimal accuracy.

The alternative formulas for  $\sin \emptyset$ ,  $\cot \alpha_{1-2}$  and  $\cot \alpha_{2-1}$ , which are given below are recommended for short lines. (They may also be used for long lines).

$$\sin \sqrt{f} = \pm \sqrt{(\sin L \cos \beta_2)^2 + (\sin \beta_2 - \beta_1) + 2\cos \beta_2 \sin \beta_1 \sin^2 \frac{L}{2})^2}$$

$$\cot \alpha_{1-2} = \frac{\left[\frac{\sin(\beta_2 - \beta_1) + \cos \beta_2 \sin \beta_1 (1 - \cos \lambda)}{\cos \beta_2 \sin \lambda}\right]}{\cos \beta_2 \sin \lambda}$$

$$\cot \alpha_{2-1} = \frac{\left[\frac{\sin(\beta_2 - \beta_1) - \cos \beta_1 \sin \beta_2 (1 - \cos \lambda)}{\cos \beta_1 \sin \lambda}\right]}{\cos \beta_1 \sin \lambda}$$
where  $\beta_2 - \beta_1 = \beta_2 - \beta_1 + \left\{n(\sin 2\beta_1 - \sin 2\beta_2) - \frac{N^2}{2}(\sinh \beta_1 - \sinh \beta_2) + \frac{n^3}{3}(\sinh \beta_1 - \sinh \beta_2)\right\}$ 
and
$$n = \frac{1 - \sqrt{1 - e^2}}{1 + \sqrt{1 - e^2}}$$

(B<sub>1</sub> and F<sub>2</sub> as previously defined are the recommend latitudes of points 1 and ?, respectively). Oso the quadrant oriterial on page 28.

### STOTIC V.

TATULAR AUTO ELECTRONIC COMPUTER SETUROD FOR SOLUTION OF DIRECT GROUNDING FROBLES, BASED OF BLOCK OF TAPLE

The formulas given below for the coldien of the direct pecietic problem are intended intended in partity for electronic computer programming. Forever, they may also be used for computation by means of deak calculators.

The commutation form for this method is as follows:

Given: B<sub>1</sub>, L<sub>1</sub> = prographic latitude and longitude, respective of any point 1.

Required:  $B_2$ ,  $L_2$  and  $A_{2-1}$ . (Latitudes and longitudes considered (+) north and east, (-) south and west).

$$\int_{C} = \left[ \int_{S} \right] \\
+ a_{1} \left[ -\frac{e^{12}}{2} \sin \beta_{S} \right] \\
+ m_{1} \left[ -\frac{e^{12}}{4} \int_{S} + \frac{e^{12}}{4} \sin \beta_{S} \cos \beta_{S} \right] \\
+ a_{1}^{2} \left[ \frac{6e^{14}}{8} \sin \beta_{S} \cos \beta_{S} \right] \\
+ m_{1}^{2} \left[ \frac{11e^{14}}{64} \int_{S} -\frac{13e^{14}}{64} \sin \beta_{S} \cos \beta_{S} - \frac{e^{14}}{8} \int_{S} \cos^{2} \beta_{S} \right] \\
+ \frac{5e^{14}}{32} \sin \beta_{S} \cos \beta_{S} - \frac{5e^{14}}{8} \sin \beta_{S} \cos^{2} \beta_{S} \right] \\
+ a_{1}m_{1} \left[ \frac{3e^{14}}{8} \sin \beta_{S} + \frac{e^{14}}{4} \int_{S} \cos \beta_{S} - \frac{5e^{14}}{8} \sin \beta_{S} \cos^{2} \beta_{S} \right] \\
+ m_{1} \left[ \frac{3f^{2}}{4} \int_{S} \sin \beta_{S} - \frac{3f^{2}}{4} \sin \beta_{S} \cos \beta_{S} \right] \\
+ m_{1} \left[ \frac{3f^{2}}{4} \int_{S} \sin \beta_{S} - \frac{3f^{2}}{4} \sin \beta_{S} \cos \beta_{S} \right]$$

where: a<sub>0</sub>, b<sub>0</sub> = semi-major and semi-minor axes, respectively, of

f = spheroidal flattening =  $\left(1 - \frac{b_0}{a_1}\right)$ 

 $e^{i2}$  = second eccentricity squared =  $(a_0^2 - b_0^2) = b_0^2$ 

 $\beta$  = number of seconds in one radians = 206,264.80625  $\beta_{8}$  = (S • b<sub>0</sub>) radians  $\beta_{8}$  = (tan B) (1-f) when  $\beta_{8} \le h_{5}^{0}$  or  $\cot \beta_{8} = \frac{(\cot \beta_{8})}{(1-f)}$  when

 $\cos \beta_0 = \cos \beta_1 \sin \alpha_{1-2}$   $g = \cos \beta_1 \cos \alpha_{1-2}$   $g_1 = (1 + \frac{e^{2}}{2} \sin^2 \beta_1) (\sin^2 \beta_1 \cos \beta_2 + g \sin \beta_1 \sin \beta_2)$   $m_1 = (1 + \frac{e^{2}}{2} \sin^2 \beta_1) (1 - \cos^2 \beta_0)$ 

 $\sin \beta_2 = \sin \beta_1 \cos \beta_0 + \varepsilon \sin \beta_0$   $\cos \beta_2 = + \sqrt{(\cos \beta_0)^2 + (\sin \beta_1 \sin \beta_0 - g \cos \beta_0)^2}$ 

The quantity under the radical and its root must be computed by showinto point notation to strain cost to full accuracy at large absolute tan  $\beta_2 = \frac{\sin \beta_2}{\cos \beta_2}$  or  $\cot \beta_2 = \frac{\cos \beta_2}{\sin \beta_2}$ , whichever has the

smaller absolute value.

Obtain tan (or cct) of  $B_2$  from its relation to tan (or cct) of  $\beta_2$ Obtain  $B_2$ , which ranges from  $-90^{\circ}$  through  $+90^{\circ}$  and takes the sign of its tan (or cot).

 $\cot \alpha_{2-1} = (\cos \alpha_{1-2} \cos \beta_0 - \tan \beta_1 \sin \beta_0) : \sin \alpha_{1-2}$ 

When  $|\cot \phi_{2-1}| > 1$ , divide result into 1 to obtain  $\tan \phi_{2-1}$  instead. (Crit these last two lines for meridional eres)

	Quadrant of $\alpha_{2-1}$			
If (08 04 1-2 \$ 180°)	and cot (or tan) of $\alpha_{2-1}$ is (+) or (-), $\alpha_{2-1}$ is in Quad. III or IV, respectively.			
If (180% × 1-2 < 360°)	and cot (or tan) of $\alpha_{2-1}$ is (+) or (-), $\alpha_{2-1}$ is in Quad. I or II, respectively.			

For meridional arcs, enter the above table with the sign of the numerator of  $\cot \alpha_{2-1}$ , and reference angle  $0^{\circ}$ .

 $\cot \lambda = (\cot \mathcal{J}_0 \cos \beta_1 - \cos \alpha_{1-2} \sin \beta_1) + \sin \alpha_{1-2}$  When  $|\cot \lambda| > 1$ , divide result into 1 to obtain ten  $\lambda$  instead.

(thit fight left two lines for mericional ares)

f	Quadrant and Sign of	<i>λ</i>
	When $\phi_0$ is in Quad.I or II (180° included)	When on is in Quad. III
and (0°≤ α′ <sub>1-2</sub> ≤180°)	then if cot (or tan) of Ais (+) or (-) A is in Quad. For II,	_
and (180% cx <sub>1-2</sub> < 360°)	then if cot (or tan) of A is (+) or (-) the assoc. angle is in Quad. III or IV, respectively, and A is obtained by subtracting 360°	then if cot (or tan) of $\lambda$ is (+) or (-) the assoc. andle is in Quad. I or II, respectively, and $\lambda$ is obtained by subtracting $360^{\circ}$

For meridional arcs, enter the above table with the sign of the numerator of  $\cot \lambda$ , and reference angle  $0^{\circ}$ .

 $L_2 = L_1 + L$ [If  $|L_2| > 180^\circ$ , modify  $L_2$  by adding or subtracting  $360^\circ$  according to whether it is initially negative or positive.

## 12. Extension of Series of the Direct Geodetic Problem For Greater Accuracy

The  $e^{i6}$  term of the preceding  $J_0$  series has been derived and numerically maximized in order to determine the error in the  $J_0$  series which would result from an omission of the  $e^{i6}$  term. This maximum value is given in Appendix I, Eart A.

Likewise, the  $(\frac{L-\lambda}{P\cos\beta_0})$  series has been extended to include the  $f^3$  term. A maximum numerical value is given in Appendix 1, 1900.

The errors which the omission of the e<sup>16</sup> term of  $\not D_0$  and the r<sup>3</sup> term of  $\left(\frac{L-\lambda}{P\cos D_0}\right)$  could finally produce in  $B_2$ ,  $L_2$  and  $rac{1}{2}$  are also shown in Appendix I, Part A.

The  $e^{i\delta}$  term of the  $\not\!\! D_0$  series is as follows:

$$\begin{bmatrix} e^{15} \text{ term of } \sqrt[6]{5} = a_1^3 \left[ -\frac{29}{24} e^{15} \sin \sqrt[6]{5} \cos^2 \sqrt[6]{5} + \frac{5}{24} e^{15} \sin \sqrt[6]{5} \right] \\ + a_1^2 m_1 \left[ \frac{5}{32} e^{15} \sqrt[6]{5} - \frac{5}{8} e^{15} \sqrt[6]{5} \cos^2 \sqrt[6]{5} - \frac{13}{32} e^{16} \sin \sqrt[6]{5} \cos \sqrt[6]{5} \right] \\ + \frac{29}{16} e^{16} \sin \sqrt[6]{5} \cos^3 \sqrt[6]{5} \end{bmatrix}$$

+ 
$$a_1 m_1^2 \left[ -\frac{39}{64} e^{-i6} f_s \cos f_s + \frac{5}{8} e^{-i6} f_s \cos 3 f_s + \frac{e^{-i6}}{32} f_s^2 \sin f_s \cos f_s + \frac{5}{8} e^{-i6} f_s \cos 3 f_s + \frac{e^{-i6}}{32} f_s^2 \sin f_s \cos f_s + \frac{e^{-i6}}{64} e^{-i6} \sin f_s \cos f_s + \frac{e^{-i6}}{32} f_s^2 \cos f_s^2 + \frac{e^{-i6}}{64} e^{-i6} \sin f_s \cos f_s + \frac{29}{32} e^{-i6} f_s \cos f_s + \frac{35}{128} e^{-i6} f_s \cos f_s^2 - \frac{5}{32} e^{-i6} f_s \cos f_s^2 + \frac{15}{256} e^{-i6} \sin f_s \cos f_s + \frac{15}{256} e^{-i6} \sin f_s \cos f_s + \frac{15}{256} e^{-i6} \sin f_s \cos f_s + \frac{59}{192} e^{-i6} \sin f_s \cos f_s^2 + \frac{29}{192} e^{-i6} \sin f_s \cos f_s^2 \right] + a_1^2 t_1 \left[ \frac{3}{16} e^{-i6} \sin f_s \cos f_s - \frac{3}{16} e^{-i6} \sin f_s \cos f_s^2 \right] + a_1 m_1 t_1 \left[ \frac{3}{16} e^{-i6} f_s \cos f_s - \frac{3}{16} e^{-i6} \sin f_s \cos f_s^2 \right] + m_1 t_1^2 \left[ -\frac{e^{-i6}}{16} f_s + \frac{e^{-i6}}{16} \sin f_s \cos f_s^2 \right] + m_1^2 t_1 \left[ \frac{e^{-i6}}{128} f_s - \frac{3}{32} e^{-i6} f_s \cos f_s + \frac{3}{64} e^{-i6} \sin f_s \cos f_s^2 \right] + \frac{1}{128} e^{-i6} \sin f_s \cos f_s^2 + \frac{3}{64} e^{-i6} \sin f_s \cos f_s^2 \right]$$

In the above equation  $t_1 = \sin^2 \beta_1$  and all other quantities are the same as defined on page 33.

As  $\vec{p}_s$  approaches  $180^\circ$ , each of the two terms containing  $\csc\vec{p}_s$  in the series above approaches infinity. However, they may be combined into a single finite term as follows:

The series for  $(\frac{L-\lambda}{\rho_{\cos\rho_0}})$  extended through the f<sup>3</sup> term is as

follows:

$$\frac{(L-\lambda)}{P\cos\beta} = \left[-f\vec{\phi}_{S}\right] + a_{1}\left[\left(\frac{3}{2}f^{2} + 2f^{3}\right)\sin\phi_{S}\right] \\
+ m_{1}\left[\left(\frac{3}{4}f^{2} + f^{3}\right)\vec{\phi}_{S} + \left(-\frac{3}{4}f^{2} - f^{3}\right)\sin\phi_{S}\cos\phi_{S}\right] \\
+ a_{1}^{2}\left[-4f^{3}\sin\phi_{S}\cos\phi_{S}\right] \\
+ a_{1}m_{1}\left[-\frac{5}{2}f^{3}\sin\phi_{S} - \frac{3}{2}f^{3}\phi_{S}\cos\phi_{S}\right] \\
+ 4f^{3}\sin\phi_{S}\cos^{2}\phi_{S} \\
+ m_{1}^{2}\left[-\frac{9}{8}f^{3}\phi_{S} + \frac{3}{4}f^{3}\phi_{S}\cos\phi_{S}\right] \\
+ m_{1}^{2}\left[-\frac{9}{8}f^{3}\phi_{S} + \frac{3}{4}f^{3}\phi_{S}\cos\phi_{S}\right]$$

APPENDIX I.

FrRT A

SUCARY TEFER OF BRACES IN DISTANDE & AZEATM IN FABULAR AND LEBATRONIO COMPUTER LATIO FOR MONITORINE SOLITION OF GEORETIC DIVERSE

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The errors above result from the omission of the  $f^2$  terms from the  $\frac{S}{D_0}$  and  $\frac{A-L}{9c}$  series. Computations were done for high latitudes and are based on the International Spheroid.

\* The magnitude of the error in  $\alpha$  for lines of approximately 15 miles clearly indicates the necessity of the f3 term in the  $\lambda$ -1 series for accuracy of three decimal places of seconds of arc.

E TAPA

SUMMENT TABLE OF ERRORS IN POSITION & AZIMUTH IN TARLINAR AND FLEGTRONIC CONTINUE NOT NOT TRACT CONTINUE TO THE TABLE TO THE TOTAL TO THE TABLE TABLE TO THE TABLE TABLE TO THE TABLE T

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 $\frac{1-\lambda}{1-\lambda}$  series. Computations were done for bigh leftinges and are based on the Invaritional schemics. The errors above result from the emission of the e<sup>16</sup> term from the  $Z_0$  series and the f3 term of

### APPENDIX II. BIPLIOGRAPHY

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